# Quantum Harmonic Oscillator

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#### **Abstract**

A detailed derivation of quantum harmonic oscillator solution.

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We start from the Schrödinger equation

$$
H\psi(x,t) = i\hbar \frac{\partial}{\partial t}\psi(x,t).
$$
 (1)

The eigenstates of energy satisfy the following equation:

$$
H\psi(x,t) = E\psi(x,t) = i\hbar \frac{\partial}{\partial t}\psi(x,t).
$$
\n(2)

The differential equation is separable with the solution:

$$
\psi(x,t) = e^{\frac{i}{\hbar}Et}\psi(x). \tag{3}
$$

The classical Hamiltonian for particle of mass *m* and in a quadratic potential angular frequency *ω* reads

$$
H = \frac{p^2}{2m} + \frac{1}{2}m\omega^2 x^2,
$$
\n(4)

where  $\omega$  is the natural frequency of the oscillator. As we move from the classical system to the quantum system, we upgrade the position and momentum parameters to quantum operators:

$$
x \to \hat{x}, \text{ and } p \to \hat{p}, \tag{5}
$$

where we added the "hat" to remind ourselves that these are operators. The quantum Hamiltonian becomes:

$$
H = \frac{\hat{p}^2}{2m} + \frac{1}{2}m\omega^2 \hat{x}^2.
$$
 (6)

There are two main methods to calculate the energy eigenstates.

### <span id="page-0-0"></span>**1 The hard way**

We first follow the bruteforce method. We have a second order differential equation, and we bite the bullet and sit down to solve it. We can use the coordinate basis where  $\hat{x}$  and  $\hat{p}$  have the following representations:

$$
\hat{x} = x, \text{ and } \hat{p} = -i\hbar \frac{d}{dx}.
$$
\n<sup>(7)</sup>

Therefore, the coordianate part of the Schrödinger equation becomes:

$$
\left(-\frac{\hbar^2}{2m}\frac{d^2}{dx^2} + \frac{1}{2}m\omega^2 x^2\right)\psi(x) = E\psi(x).
$$
\n(8)

We can define a couple of dimensionless quantities  $\chi = x\sqrt{\frac{m\omega}{\hbar}}$  and  $\epsilon = \frac{2E}{\hbar\omega}$  to get:

<span id="page-1-0"></span>
$$
\left(-\frac{d^2}{dx^2} + \chi^2 - \epsilon\right)\psi(x) = 0.\tag{9}
$$

We should first try to understand the asymptotic solution where  $\chi^2 \gg \epsilon$ .

$$
\left(-\frac{d^2}{dx^2} + \chi^2\right)\psi(x) \simeq 0.\tag{10}
$$

This equation has a special solution called parabolic cylinder functions. However, since we are looking for the asymptotic solutions, we can make an educated guess of the form  $e^{-\alpha \chi^2}$  and plug it in to find that  $\alpha = 1/2$ solves the differential equation at the first order. Or, we can try to split the  $-\frac{d^2}{dx^2} + \chi^2$  operator into two first order operators and drop a small term in the large  $\chi$  limit:

$$
\left(-\frac{d^2}{dx^2} + \chi^2\right)\psi(x) \simeq \left(-\frac{d}{dx} + \chi\right)\left(\frac{d}{dx} + \chi\right)\psi(x) \simeq 0. \tag{11}
$$

We then combine the asymptotic solution with a yet-unknown function and propose a solution of the following form:

<span id="page-1-3"></span>
$$
\psi(\chi) = e^{-\frac{\chi^2}{2}}h(\chi),\tag{12}
$$

upto the normalization constant, which we will calculate later. Plugging this back into Eq. [\(9\)](#page-1-0), we get the following second order differential equation.

$$
\left(-\frac{d^2}{dx^2} + 2\chi \frac{d}{dx} + 1 - \epsilon\right)h(\chi) = 0.
$$
\n(13)

We can now try a power series of the form

<span id="page-1-1"></span>
$$
h(\chi) = \sum_{j=0}^{\infty} c_j \chi^j.
$$
 (14)

Inserting this back in, we get:

$$
-\sum_{j=0}^{\infty} j(j-1)c_j \chi^{j-2} + \sum_{j=0}^{\infty} (2j+1-\epsilon)c_j \chi^j = 0.
$$
 (15)

Since the first two terms in the first summation vanish due to the  $j(j-1)$  coefficient, we can start the first summation index, *j*, from 2, and redefine *j* as  $j + 2$  and pull the starting point back to 0:

$$
-\sum_{j=0}^{\infty} (j+2)(j+1)c_{j+2}\chi^j + \sum_{j=0}^{\infty} (2j+1-\epsilon)c_j\chi^j = \sum_{j=0}^{\infty} \left( -(j+2)(j+1)c_{j+2} + (2j+1-\epsilon)c_j \right)\chi^j = 0. \tag{16}
$$

In order to set this to zero we need to have the recurrence equation:

<span id="page-1-2"></span>
$$
c_{j+2} = \frac{2j+1-\epsilon}{(j+2)(j+1)}c_j.
$$
\n(17)

Note that this is problematic because the coefficients are not decaying fast enough. In fact, this relation implies that  $h(\chi) \propto e^{\chi^2}$ , and even the prefactor  $e^{-\chi^2/2}$  will not be decaying fast enough to make the wave-function normalizable. The only way out of this is to truncate the series at some point. Remember that the only knob we have is  $\epsilon$ , and we can set it an integer value such that when  $2j + 1 = \epsilon$ , the series terminates. This is a profound finding because the physicality of the solution requires the quantization of the energy. Going back to the original parameters,  $E = \frac{\epsilon \hbar \omega}{2}$ , we can write the energy as:

<span id="page-2-3"></span>
$$
E = \hbar\omega(n + \frac{1}{2}).\tag{18}
$$

There is another subtle problem: note that the recurrence formula relates  $c_0$  to  $c_2$ ,  $c_2$  to  $c_4$  so and so forth, and  $c_1$  to  $c_3$ ,  $c_3$  to  $c_5$ . In other words, the only free coefficients are  $c_0$  and  $c_1$ . As we discussed earlier, we can truncate the series by selecting  $\epsilon$  appropriately. However, we have only one degree of freedom in  $\epsilon$ , and we can't use that to terminate both of the odd and even series at the same time. That means only one of the coefficients *c*<sup>0</sup> and *c*<sup>1</sup> can be non-zero at the same time. This is also expected from the parity symmetry of the Hamiltonian: it stays invariant under  $x \to -x$ , which implies that solutions should stay invariant up to the sign. Therefore, odd and even powers of  $\chi$  cannot mix in the energy eigenstates.

Since the series in Eq. [\(14\)](#page-1-1) will terminate at  $j = n/2$  due to the recurrence relation in Eq. [\(17\)](#page-1-2), it is convenient to redefine the summation index  $s = \frac{n}{2} - j$ , and rewrite the solution as a sum of finite number of terms:

<span id="page-2-0"></span>
$$
h_n(\chi) = \sum_{s=0}^{\lfloor n/2 \rfloor} (-1)^s \frac{n!}{(n-2s)!s!} (2x)^{n-2s},\tag{19}
$$

where  $(-1)^s$  originates from flipping the sign of the numerator in Eq. [\(17\)](#page-1-2), and powers of 2 originate from  $n/2$ 's in the denominator. These are Hermite polynomial, which can be written explicitly as

$$
h_0(\chi) = 1\nh_1(\chi) = 2x\nh_2(\chi) = 4x^2 - 2\nh_3(\chi) = 8x^3 - 12x\nh_4(\chi) = 16x^4 - 48x^2 + 12
$$
\n(20)\n
$$
\vdots
$$
\n(21)

Now we have to deal with the normalization of the wavefunction in Eq. [\(12\)](#page-1-3). There is a very elegant way of doing this via the generating function. Let's multiply Eq. [\(19\)](#page-2-0) with  $\frac{t^n}{n!}$  $\frac{t^n}{n!}$  and sum over *n*:

<span id="page-2-2"></span>
$$
g(\chi, t) \equiv \sum_{n=0}^{\infty} \frac{t^n}{n!} h_n(\chi) = \sum_{n=0}^{\infty} \sum_{s=0}^{\lfloor n/2 \rfloor} (-1)^s \frac{t^n}{(n-2s)!s!} (2\chi)^{n-2s} = \sum_{n=0}^{\infty} \sum_{s=0}^{\infty} (-1)^s \frac{t^n}{(n-2s)!s!} (2\chi)^{n-2s},
$$
 (22)

where we extended the summation upper limit since we will negative factorials give negative infinities killing all the terms for  $s > n/2$ . Now define  $n - 2s = m$  and do some shuffling:

$$
g(\chi, t) = \sum_{m=0}^{\infty} \sum_{s=0}^{\infty} (-1)^s \frac{t^{m+2s}}{m!s!} (2\chi)^m = \sum_{s=0}^{\infty} \frac{(-t^2)^s}{s!} \sum_{m=0}^{\infty} \frac{(2t\chi)^m}{m!} = e^{-t^2 + 2t\chi}.
$$
 (23)

Now it becomes an easy task to compute the normalization factor. Consider the following:

<span id="page-2-1"></span>
$$
\int_{-\infty}^{\infty} d\chi e^{-\chi^2} g(\chi, t) g(\chi, q) = \int_{-\infty}^{\infty} d\chi e^{-\chi^2 - t^2 + 2t\chi - q^2 + 2q\chi} = \int_{-\infty}^{\infty} d\chi e^{-(\chi - t - q)^2 + 2qt} = \sqrt{\pi} e^{2qt} = \sqrt{\pi} \sum_{n=0}^{\infty} \frac{(2qt)^n}{n!} (2qt)^n
$$

and evaluate the integral in the series expansion:

<span id="page-3-0"></span>
$$
\int_{-\infty}^{\infty} d\chi e^{-\chi^2} g(\chi, t) g(\chi, q) = \int_{-\infty}^{\infty} d\chi e^{-\chi^2} \sum_{n=0}^{\infty} \frac{t^n}{n!} h_n(\chi) \sum_{m=0}^{\infty} \frac{q^m}{m!} h_m(\chi) = \sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \frac{t^n}{n!} \frac{q^m}{m!} \int_{-\infty}^{\infty} d\chi e^{-\chi^2} h_n(\chi) h_m(\chi)
$$

$$
= \sum_{n=0}^{\infty} \frac{(qt)^n}{(n!)^2} \int_{-\infty}^{\infty} d\chi e^{-\chi^2} h_n(\chi) h_n(\chi) + \sum_{n=0}^{\infty} \sum_{m=0, m \neq n}^{\infty} \frac{t^n}{n!} \frac{q^m}{m!} \int_{-\infty}^{\infty} d\chi e^{-\chi^2} h_n(\chi) h_m(\chi) \delta(\chi) d\chi
$$

By comparing the coefficients of *qt* terms in Eqs. [\(24\)](#page-2-1) and [\(25\)](#page-3-0), we first see that the cross terms should vanish. We also get the normalization constant:

$$
\int_{-\infty}^{\infty} d\chi e^{-\chi^2} h_n(\chi) h_m(\chi) = 2^n \sqrt{\pi} n! \delta_{n,m}.
$$
\n(26)

The orhogonality of the eigenfunctions is not a coincidence since the differential operator we are dealing with can be transformed into a self-adjoint form, and the orthogonality is guaranteed due to the Sturm-Liouville theor[y\[1\].](#page-5-0) Putting it all together we have the normalized energy eigenstates:

<span id="page-3-3"></span>
$$
\psi_n(x) = \left(\frac{m\omega}{\pi\hbar}\right)^{\frac{1}{4}} \frac{1}{\sqrt{2^n n!}} h_n\left(\sqrt{\frac{m\omega}{\hbar}}x\right) e^{-\frac{m\omega x^2}{2\hbar}}, \quad n = 0, 1, 2, \dots \quad . \tag{27}
$$

It is operationally more practical to combine  $\hat{x}$  and  $\hat{p}$  operators into raising and lowering ladder operators. The harder method is based on the recurrence relations of the Hermite polynomials. Taking the derivative of the equality in Eq. [\(22\)](#page-2-2) with respect to  $\chi$ , we get:

$$
\frac{\partial}{\partial \chi} g(\chi, t) = 2te^{-t^2 + 2t\chi} = 2 \sum_{n=0}^{\infty} \frac{t^{n+1}}{n!} h_n(\chi) = 2 \sum_{m=1}^{\infty} \frac{t^m}{(m-1)!} h_{m-1}(\chi) = 2 \sum_{m=1}^{\infty} \frac{mt^m}{m!} h_{m-1}(\chi) = 2 \sum_{n=0}^{\infty} \frac{nt^n}{n!} h_{n-1}(\chi)
$$
\n
$$
= \sum_{n=0}^{\infty} \frac{t^n}{n!} h_n'(\chi),
$$
\n(28)

where we first defined  $m = n + 1$ , and then relabeled m as n. We also added the vanishing  $n = 0$  term in the summation to make the sum start from 0. Matching the coefficients of  $t^n$  terms, we get the first recurrence relation of the Hermite functions:

<span id="page-3-1"></span>
$$
2nh_{n-1}(\chi) = h'_n(\chi). \tag{29}
$$

Let's try taking the derivative with respect to *t* to get:

$$
\frac{\partial}{\partial t}g(\chi,t) = (-2t+2\chi)e^{-t^2+2t\chi} = -2\sum_{n=0}^{\infty} \frac{t^{n+1}}{n!}h_n(\chi) + 2\sum_{n=0}^{\infty} \frac{t^n}{n!}\chi h_n(\chi) = \sum_{n=0}^{\infty} \frac{t^n}{n!} \{2\chi h_n(\chi) - 2nh_{n-1}(\chi)\}
$$
\n
$$
= \sum_{n=0}^{\infty} n \frac{t^{n-1}}{n!}h_n(\chi) = \sum_{n=0}^{\infty} \frac{t^n}{n!}h_{n+1}(\chi). \tag{30}
$$

Matching the coefficients of  $t^n$  terms, we get the second recurrence relation of the Hermite functions:

<span id="page-3-2"></span>
$$
h_{n+1}(\chi) = 2\chi h_n(\chi) - 2nh_{n-1}(\chi). \tag{31}
$$

We can combine Eqs. [\(29\)](#page-3-1) and [\(31\)](#page-3-2) to get another flavor:

<span id="page-3-4"></span>
$$
h_{n+1}(\chi) = \left(2\chi - \frac{d}{d\chi}\right)h_n(\chi). \tag{32}
$$

Now consider the following operator acting on  $\psi_n(\chi)$  as it is defined Eq. [\(27\)](#page-3-3):

$$
\frac{1}{\sqrt{2}} \left( \chi - \frac{d}{d\chi} \right) \psi_n(\chi) = \frac{1}{\sqrt{2}} \left( \frac{m\omega}{\pi\hbar} \right)^{\frac{1}{4}} \frac{1}{\sqrt{2^n n!}} \left( \chi - \frac{d}{d\chi} \right) \left( h_n(\chi) e^{-\frac{\chi^2}{2}} \right)
$$
\n
$$
= \left( \frac{m\omega}{\pi\hbar} \right)^{\frac{1}{4}} \frac{\sqrt{n+1}}{\sqrt{2^{n+1}(n+1)!}} e^{-\frac{\chi^2}{2}} \left( 2\chi - \frac{d}{d\chi} \right) h_n(\chi) = \sqrt{n+1} \psi_{n+1}(\chi), \quad (33)
$$

where we used Eqs. [\(32\)](#page-3-4).

Now consider another operator acting on  $\psi_n(\chi)$ :

$$
\frac{1}{\sqrt{2}}\left(\chi + \frac{d}{d\chi}\right)\psi_n(\chi) = \frac{1}{\sqrt{2}}\left(\frac{m\omega}{\pi\hbar}\right)^{\frac{1}{4}}\frac{1}{\sqrt{2^n n!}}\left(\chi + \frac{d}{d\chi}\right)\left(h_n(\chi)e^{-\frac{\chi^2}{2}}\right) = \left(\frac{m\omega}{\pi\hbar}\right)^{\frac{1}{4}}\frac{1}{\sqrt{2^{n+1}n!}}e^{-\frac{\chi^2}{2}}h'_n(\chi)
$$
\n
$$
= \left(\frac{m\omega}{\pi\hbar}\right)^{\frac{1}{4}}\frac{1}{2\sqrt{n}\sqrt{2^{n-1}(n-1)!}}e^{-\frac{\chi^2}{2}}2nh_{n-1}(\chi) = \sqrt{n}\psi_{n-1}(\chi),\tag{34}
$$

where we used Eqs. [\(29\)](#page-3-1). The operators  $\frac{1}{\sqrt{2}}$  $\frac{1}{2}\left(\chi \pm \frac{d}{d\chi}\right)$  can be written interms of *x* and  $\hat{p}$  and they will be called *a* and  $a^{\dagger}$ , and that would be how one solves the harmonic oscillator the hard way. Now let's look into the method of operators.

## <span id="page-4-0"></span>**2 The operational way**

We can define the ladder operators as follows:

$$
a = \frac{1}{\sqrt{2m\hbar\omega}}(m\omega x + ip), \quad a^{\dagger} = \frac{1}{\sqrt{2m\hbar\omega}}(m\omega x - ip) \iff \hat{x} = \sqrt{\frac{\hbar}{2m\omega}}(a+a^{\dagger}), \quad \hat{p} = -i\sqrt{\frac{m\hbar\omega}{2}}(a-a^{\dagger}). \tag{35}
$$

The commutation relation  $[x, p] = i\hbar$  turns to

$$
[a, a^{\dagger}] = 1. \tag{36}
$$

The Hamiltonian becomes

<span id="page-4-1"></span>
$$
H \equiv \hbar\omega(a^\dagger a + \frac{1}{2}).\tag{37}
$$

Comparing Eq. [\(37\)](#page-4-1) with Eq. [\(18\)](#page-2-3) we can associate  $a^{\dagger}a$  with number operator:

$$
N = a^{\dagger} a,\tag{38}
$$

which returns the state number:

$$
N|n\rangle = n|n\rangle. \tag{39}
$$

Let's now figure out how *a* and  $a^{\dagger}$  act on eigenstate  $|n\rangle$ . We can read the energy value by acting on the new state with *H*:

$$
Ha|n\rangle = (aH + [H, a])|n\rangle = (aH - a\hbar\omega)|n\rangle = \hbar\omega\left(n - 1 + \frac{1}{2}\right)a|n\rangle,
$$
\n(40)

which shows that the state  $a|n\rangle$  can be indexed as  $n-1$ , i.e.,  $a|n\rangle = c|n-1\rangle$  where *c* is the normalization constant. The overall coefficient *c* can be calculated as

$$
|a|n\rangle|^2 = \langle n|a^\dagger a|n\rangle = n\langle n|n\rangle = n = |c|^2 \implies c = \sqrt{n}.\tag{41}
$$

Therefore the lowering operator *a* does the following:

<span id="page-4-2"></span>
$$
a|n\rangle = \sqrt{n}|n-1\rangle.
$$
 (42)

As a consequence of Eq. [\(42\)](#page-4-2), we see that the ground state, |0⟩, will be annihilated by the operator *a*

$$
a|0\rangle = 0.\t\t(43)
$$

Let's repeat for  $a^{\dagger}$ :

$$
Ha^{\dagger}|n\rangle = \left(a^{\dagger}H + [H, a^{\dagger}]\right)|n\rangle = \left(a^{\dagger}H + a^{\dagger}\hbar\omega\right)|n\rangle = \hbar\omega\left(n+1+\frac{1}{2}\right)a|n\rangle,\tag{44}
$$

which shows that the state  $a^{\dagger} | n \rangle$  can be indexed as  $n + 1$ , i.e.,  $a^{\dagger} | n \rangle = d | n + 1 \rangle$  where *d* is the normalization constant. The overall coefficient *d* can be calculated as

$$
|a^{\dagger}|n\rangle|^2 = \langle n|aa^{\dagger}|n\rangle = \langle n|a^{\dagger}a + [a,a^{\dagger}]|n\rangle = (n+1)\langle n|n\rangle = n = |d|^2 \implies d = \sqrt{n+1}.
$$
 (45)

Therefore the lowering operator  $a^{\dagger}$  does the following:

$$
a^{\dagger}|n\rangle = \sqrt{n+1}|n+1\rangle.
$$
 (46)

By recursively applying  $a^{\dagger}$  on  $|0\rangle$  we can get the *n*-th energy eigenstate,  $|n\rangle$ :

$$
|n\rangle = \frac{(a^{\dagger})^n}{\sqrt{n!}}|0\rangle. \tag{47}
$$

Let's take a look at certain expectation values. We can immediately see that the expected values of *x* and  $\hat{p} = i\hbar \frac{d}{dx}$  vanish since the integrands of  $\langle \psi_n | x | \psi_n \rangle$  and  $\langle \psi_n | \hat{p} | \psi_n \rangle$  are odd and the integration range is symmetric around the origin. Equivalently we can do the computation using the ladder operators:

$$
\langle \hat{x} \rangle_n = \langle n | \hat{x} | n \rangle = \sqrt{\frac{\hbar}{2m\omega}} \langle n | (a + a^\dagger) | n \rangle = 0.
$$
 (48)

Similarly for *p*, we have:

$$
\langle \hat{p} \rangle_n = \langle n | \hat{p} | n \rangle = -i \sqrt{\frac{m \hbar \omega}{2}} \langle n | a - a^\dagger | n \rangle = 0. \tag{49}
$$

Now consider the quadratic operators:

$$
\langle \hat{x}^2 \rangle_n = \langle n | \hat{x}^2 | n \rangle = \frac{\hbar}{2m\omega} \langle n | (a + a^\dagger)^2 | n \rangle = \frac{\hbar}{2m\omega} \langle n | (2a^\dagger a + [a, a^\dagger]) | n \rangle = \frac{\hbar}{2m\omega} (2n + 1). \tag{50}
$$

Similarly for *p*, we have:

$$
\langle \hat{p}^2 \rangle_n = \langle n | \hat{p}^2 | n \rangle = -\frac{m \hbar \omega}{2} \langle n | (a - a^\dagger)^2 | n \rangle = \frac{m \hbar \omega}{2} \langle n | 2a^\dagger a + [a, a^\dagger] | n \rangle = \frac{\hbar}{2m\omega} (2n + 1). \tag{51}
$$

The uncertainity in *x* and *p* for state *n* are given as:

$$
\langle (\Delta x)^2 \rangle_n = \langle \hat{x}^2 \rangle_n - (\langle \hat{x} \rangle_n)^2 = \frac{\hbar}{2m\omega} (2n+1), \tag{52}
$$

and

$$
\langle (\Delta p)^2 \rangle_n = \langle \hat{p}^2 \rangle_n - (\langle \hat{p} \rangle_n)^2 = \frac{\hbar m \omega}{2} (2n + 1).
$$
 (53)

The Heisenberg relation becomes

$$
(\Delta x)^2 (\Delta p)^2 = \frac{\hbar^2}{4} (2n+1)^2,
$$
\n(54)

which has the minimum value of  $\frac{\hbar^2}{4}$  $\frac{5^{2}}{4}$  at  $n = 0$ .

<span id="page-5-0"></span>[1] G. B. Arfken and H. J. Weber, *Mathematical Methods for Physicists, 4th edition*. Academic Press, San Diego, 1995, pp. 537–547.